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A SOLUTION OF KEPLER'S PROBLEM FOR PLANETARY ORBITS OF HIGH ECCENTRICITY.

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The solution of Kepler's Problem involves the transcendental equation

$$M = E - e \sin E, \tag{1}$$

in which M and E are respectively the mean and eccentric anomalies, while e is the eccentricity.

Let

$$F(E) = \frac{1}{2} [E - M - \sin(E - M)]$$

$$= \frac{1}{12} \sin^3(E - M) + \frac{3}{80} \sin^5(E - M) + \frac{5}{224} \sin^7(E - M) + \dots$$
(2)

From (1) and (2) we easily get Grunert's equation,

$$\frac{1+e}{1-e} = \frac{\cos\frac{1}{2}M\sin(E-\frac{1}{2}M) + F(E)}{\sin\frac{1}{2}M\cos(E-\frac{1}{2}M) - F(E)}.$$
 (3)

Equation (3) gives

$$\tan(E - \frac{1}{2}M) = \frac{I + e}{I - e} \tan \frac{1}{2}M - \frac{2F(E)}{(I - e)\cos \frac{1}{2}M\cos(E - \frac{1}{2}M)}.$$
 (4)

Assume that

$$\tan (E' - \frac{1}{2}M) = \frac{I + e}{I - e} \tan \frac{1}{2}M.$$
 (5)

Substituting in (4) according to (5) and reducing, we obtain

$$\sin(E' - E) = \frac{\cos(E' - \frac{1}{2}M)}{(I - e)\cos\frac{1}{2}M} \cdot 2F(E)$$
 (6)

$$= k \cdot 2 F(E). \tag{7}$$

Hence for moderate eccentricities, E' - E is approximately equal to 2F(E).

Let
$$2\eta = E' - M - \sin(E' - M)$$
. (8)

Then E'-E is likewise nearly equal to 2η , and $E'-M-2\eta$ is nearly equal

to E - M. From (2) we obtain by differentiation

$$dF(E) = \left[\frac{1}{4}\sin^2(E-M) + \frac{3}{16}\sin^4(E-M) + \dots\right]\cos(E-M)d(E-M).$$
 (9)

Equation (9) shows that a small change in the value of E-M produces a much smaller change in F(E); therefore, we write with but slight error

$$2 F(E) = E' - M - 2\eta - \sin(E' - M - 2\eta).$$
 (10)

From (8), omitting terms of the seventh and higher orders, we have

$$\eta = \frac{1}{12}\sin^3(E' - M) + \frac{3}{80}\sin^5(E' - M). \tag{11}$$

Suppose that E' is changed by an amount equal to 2η ; i. e. that $dE' = 2\eta$. Differentiation of (11) and reduction of the resulting equation, neglecting powers of $\sin (E' - M)$ higher than the fifth, gives

$$d\eta = \frac{1}{24}\sin^5(E' - M)\cos(E' - M); \tag{12}$$

The last two terms of (13) have nearly the same magnitude, and opposite signs. We therefore write $\eta - d\eta = \frac{1}{12} \sin^3(E' - M)$. (14)

Now since (12) is obtained on the assumption that $dE' = 2\eta$,

$$\eta - d\eta = \frac{1}{2} [E' - M - 2\eta - \sin(E' - M - 2\eta)].$$

Hence by (10) and (14) we may write

$$F(E) = \frac{1}{12}\sin^3(E' - M). \tag{15}$$

Equations (15) and (7) give

$$\sin(E' - E) = \frac{1}{6}k \sin^3(E' - M). \tag{16}$$

Let $E'-M-\frac{1}{6}k\sin^3{(E'-M)}$ be an approximate value of E-M; then $2F\left[E'-M-\frac{1}{6}k\sin^3{(E'-M)}\right]$ will be the corresponding value of 2F(E), and by (7), $k\cdot 2F\left[E'-M-\frac{1}{6}k\sin^3{(E'-M)}\right]$ will be the corresponding value of $\sin{(E'-E)}$. Differentiation of (7) gives

$$\cos(E'-E) d(E'-E) = k \left[\frac{1}{2} \sin^2(E-M) + \frac{3}{8} \sin^4(E-M) + \dots \right] \cos(E-M) d(E-M)$$

$$= k \left[\frac{1}{2} \sin^2(E-M) + \frac{1}{8} \sin^4(E-M) + \dots \right] d(E-M)$$

$$= kv d(E-M).$$
(17)

If E-M be increased by $E'-E-\frac{1}{6}k\sin^3(E'-M)$, and the corresponding increment of E'-E be denoted by $\Delta(E'-E)$, we have

$$\Delta(E' - E) = kv \left[E' - E - \frac{1}{6}k \sin^3(E' - M) \right] \sec(E' - E). \tag{18}$$

But $E' - E + \Delta(E' - E) = k \cdot 2F \left[E' - M - \frac{1}{6}k \sin^3(E' - M) \right] + 2F(E' - E).$ (19)

Substitute from (18) in (19), and add and substract $\frac{1}{6}k \sin^3(E'-M)$ in the second member of the resulting equation; reduction gives

$$(E' - E)[1 + kv \sec(E' - E)] = [1 + kv \sec(E' - E)] \cdot \frac{1}{6}k \sin^3(E' - M) + k \cdot 2F[E' - M - \frac{1}{6}k \sin^3(E' - M)] - \frac{1}{6}k \sin^3(E' - M) + 2F(E' - E).$$
(20)

Division of (20) by the coefficient of E'-E, and subtraction of 2F(E'-E) from each side, gives

$$\sin(E'-E) = \frac{1}{6}k \sin^3(E'-M) + \frac{k \left\{ 2F[E'-M-\frac{1}{6}k\sin^2(E'-M)] - \frac{1}{6}\sin^3(E'-M) \right\}}{1 + kv \sec(E'-E)} - \frac{kv \sec(E'-E) \cdot 2F(E'-E)}{1 + kv \sec(E'-E)}$$
(21)

But we have, with sufficient accuracy,

$$2F(E'-E) = \frac{1}{6}\sin^3(E'-E) = \frac{1}{1296}k^3\sin^9(E'-M),$$

$$\frac{kv\sec(E'-E) \cdot 2F(E'-E)}{1+kv\sec(E'-E)} = \frac{k^4\sin^{11}(E'-M)}{2592(1+kv)}.$$
(22)

and

We next determine the error of (21) when its last term is neglected and the quantity $\sec (E'-E)$ is dropped from the denominator of the second fraction. The eccentricities of the orbits of all but 19 of the first 268 asteroids are below sin 15°. Of these 19, 15 lie between $\sin 15^\circ$ and $\sin 20^\circ$, and 4 are greater than $\sin 20^\circ$; the orbit of Aethra (132) has the greatest eccentricity, e being equal to $\sin 22^\circ$. On the assumption that $e = \sin 23^\circ$, we find that the maximum value of the last term of (21), as computed by (22), is 0.003 + 1. If the factor $\sin (E'-E)$ in the denominator of the second fraction of (21) be dropped, an error of 0.0002 may result. To compute e in the same denominator, we substitute for e m (in the expression given for e in (17)) the value e multiple e in e substitute for e multiple e multiple equations (17)–(21) we see that, had e multiple e multiple equations (17)–(21) we see that, had e multiple equations (18) and e multiple equations

$$E' = M - \frac{1}{6}k\sin^3(E' - M) - \frac{1}{2}\left[E' - E - \frac{1}{6}k\sin^3(E' - M)\right]$$

The greatest error in E'-E caused by using $E'-M-\frac{1}{6}k\sin^3(E'-M)$ for E-M, in finding v, is less than 0.002. Greater accuracy is unnecessary, but could be attained, since $\tau + kv$ is always close to unity, by using

$$E' - M - \frac{1}{6}k \sin^3(E' - M) \\ - \frac{1}{2}k \left\{ 2F \left[E' - M - \frac{1}{6}k \sin^3(E' - M) \right] - \frac{1}{6}\sin^3(E' - M) \right\},\,$$

or
$$E' - M - \frac{1}{12}k \sin^3(E' - M) - k F[E' - M - \frac{1}{6}k \sin^3(E' - M)].$$

Equation (21) may therefore be written

$$\sin(E'-E) = \frac{1}{6}k \sin^3(E'-M) + \frac{k \left\{2F\left[E'-M-\frac{1}{6}k\sin^3(E'-M)\right] - \frac{1}{6}\sin^3(E'-M)\right\}}{1+kv}.$$
 (23)

When $e = \sin 23^\circ$, the maximum numerical value of the last fraction of (23) is 38".6. A simple approximate value of this fraction in terms of k and $\sin (E' - M)$ may be obtained as follows:—

From (9) and (2), by putting E'-M for E-M, $\frac{1}{6}k\sin^3(E'-M)$ for d(E-M), and $1-\frac{1}{2}\sin^2(E'-M)$ for $\cos(E'-M)$, we get

$$2F[E'-M-\tfrac{1}{6}k\sin^3(E'-M)] = \tfrac{1}{6}\sin^3(E'-M) + (\tfrac{3}{40}-\tfrac{1}{12}k)\sin^5(E'-M) + (\tfrac{5}{112}-\tfrac{1}{48}k)\sin^7(E'-M). \tag{24}$$

Since kv is small, we assume that

$$\frac{1}{1+kv} = 1 - kv = 1 - \frac{1}{2}k\sin^2(E' - M). \tag{25}$$

Substitution of (24) and (25) in the last fraction of (23) gives for its value

$$\tfrac{1}{120}(9k-10k^2)\sin^5(E'-M) + \left[\tfrac{1}{336}(15k-7k^2) - \tfrac{1}{240}(9k^2-10k^3)\right]\sin^7(E'-M)$$

Since $\sin{(E'-M)}=e\sin{E'}$, neglecting powers higher than the fifth, the preceding expression becomes $\frac{1}{120}(9k-10k^2)e^5\sin^5{E'}$. Because this involves the fifth power of e, its value is small for moderate eccentricities; nevertheless, the expression is not sufficiently accurate to be employed in seven-place computations with large eccentricities. As the second member of (23) is not easy to compute, it may be simplified, and its errors tabulated. It is not difficult to show that the maximum value of k is $(1-e)^{-1}$, which is reached when v=0. The maximum of kv is near $\frac{1}{2}e^2$, which is only 0.08 when $e=\sin{23^\circ}$; kv being, therefore, always small, neglect of the denominator of the last term of (23) introduces into E'-E an error the value of which is not far from

$$\frac{1}{120} kv (9k - 10k^2) e^5 \sin^5 E'$$
, or $\frac{1}{240} (9k^2 - 10k^3) e^7 \sin^7 E'$.

When $e = \sin 23^{\circ}$ this error may amount to 3"; when $e = \sin 15^{\circ}$ it does not reach 0".2. Rejecting I + kv and writing E' - E for $\sin (E' - E)$, (23) becomes

$$E' - E = k \cdot 2F \left[E' - M - \frac{1}{6}k \sin^3(E' - M) \right]$$
 (26)

To expedite computation, two tables are needed, one giving the correction of the second member of (26), the other containing 2F(x) with the argument x. The computer having e and M given, would use (5) and (26), taking into account the tabulated correction to the value of E' - E given by (26). From the value of E thus found, the true anomaly could be obtained by one of the usual methods.